

## Notes On De Sitter Space

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de Sitter spacetime is the maximally symmetric spacetime of constant positive curva- ture. It is a solution of the vacuum Einstein equations with a positive cosmological constant. It is directly relevant for observation, in two (as fas as we know unrelated!) ways.

~~Lecture Notes on Classical de Sitter Space~~

de Sitter space is the subset  $deS = f_{hx;xi} = a^2 j_x 2M5g$ : There is an isometric copy  $H^4_q$  of hyperbolic space with  $x_0 < 0$ . The induced metric on hyperbolic space is Riemannian and on de Sitter space is Lorentzian. Thus de Sitter space is a space-time. It is a solution of

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Einstein's equations with positive cosmological constant  $\Lambda = 3a^2$  and no matter.

## ~~Notes on de Sitter space~~

In mathematical physics,  $n$ -dimensional de Sitter space is a maximally symmetric Lorentzian manifold with constant positive scalar curvature. It is the Lorentzian analogue of an  $n$ -sphere. The main application of de Sitter space is its use in general relativity, where it serves as one of the simplest mathematical models of the universe consistent with the observed accelerating expansion of the universe. More specifically, de Sitter space is the maximally symmetric vacuum solution of Einstein's field

## ~~de Sitter space — Wikipedia~~

Notes on Euclidean de Sitter space - NASA/ADS Note that de Sitter space has an initial and final conformal boundary. (Although the diagram also appears to have left and right boundaries, these are not really boundaries - at each value of  $t$  space is a sphere, so those lines are just the north and south poles of the sphere  $S^1$ .)

## ~~Notes On De Sitter Space~~

can always "unwrap" the hyperboloid by going to the covering space. Note that in  $1 + 1$  dimensions we can always switch the meaning of timelike and spacelike. Then we obtain de Sitter space  $dS_2$ , that has a closed space but no closed timelike curves. In general the topology of  $adS_n$  is  $R^{n-1} \times S^1$  and the topology of  $dS_n$  is  $S^{n-1} \times R$ , so that it is only in two dimensions that  $dS_4$

## ~~ANTI-DE SITTER SPACE~~

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## ~~7 Thermodynamics of de Sitter space — hartmanhep.net~~

There are ways to cast de Sitter space with static coordinates (see de Sitter space), so unlike other FLRW models, de Sitter space can be thought of as a static solution to Einstein's equations even though the geodesics followed by observers necessarily diverge as expected from the expansion of physical spatial dimensions.

## ~~de Sitter universe — Wikipedia~~

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de Sitter space has a number of good properties. It has been shown to be stable [1], and to possess positive energy representations [2] (see [3] for a review of further properties of anti-de Sitter space). Recently, anti-de Sitter space has appeared in a surprising new context. Maldacena [4] has conjectured that the large  $N$  limit of cer-

~~arXiv:hep-th/9805087v1 14 May 1998~~

Just for completeness, note that Anti de Sitter space is the maximally symmetric solution to Einstein's equations with negative cosmological constant. Finally a quick note: de Sitter (Anti de Sitter) space has constant positive (negative) scalar curvature and hence is non-hyperbolic (hyperbolic).

~~General Relativity: What is de Sitter space? Why does it ...~~

It was proved by K. Akutagawa [a1], Q.M. Cheng [a2] and K.G. Ramanathan that complete space-like submanifolds with parallel mean curvature vector in a de Sitter space  $S_{-p}^{n+p}(c)$  are totally umbilical (cf. also Differential geometry) if  $1) |H|^2 \leq c$ , when  $n = 2$ ;

~~De Sitter space — Encyclopedia of Mathematics~~

These lectures present an elementary discussion of some background material relevant to the problem of de Sitter quantum gravity. The first two lectures discuss the classical geometry of de Sitter space and properties of quantum field theory on de Sitter space, especially the temperature and entropy of de Sitter space. The final lecture contains a pedagogical discussion of the appearance of ...

~~[hep-th/0110007] Les Houches Lectures on De Sitter Space~~

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Notes on de Sitter space and holography 5657 Keeping the AdS/CFTcorrespondencein mind, we proceed to study the action for scalar fields in de Sitter space as a functional of boundary data. To extend this investigation to gravity, we display a family of solutions to three-dimensional (3D) gravity with a positive

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However, when de Sitter entropy is computed in a 'stretched horizon' picture, then we argue that the correct euclidean topology is a solid torus. The solid torus shrinks and degenerates into a three-hemisphere as one goes from the 'stretched horizon' to the horizon, giving the euclidean continuation of the causal diamond.

~~Notes on Euclidean de Sitter space — NASA/ADS~~

The isometry group of de Sitter space is therefore the Lorentz group  $S O (4, 1)$ . The embedding space coordinates are very useful for many calculations. In

~~(PDF) De Sitter Space and Spatial Topology~~

In de Sitter space there is a non-zero probability to pair-produce charged black holes from the vacuum [7, 8, 9, 10, 11, 12, 13]. The two-dimensional FRW regions in the interior of each black hole are produced dynamically, and so black hole nucleation can be regarded as the dynamical compactification of two extra dimensions.

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