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The emission lines in
Be stars are thought
to originate from a low-
latitude disk or ring-
like envelope (Kogure
& Leung 2007), which
in the case of BD+50
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arXiv:2009.04193

(astro-ph) [Submitted
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A Broad Absorption
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SDSS J075133.35+1
34548.3. Authors: Bo
Liu, Hongyan Zhou,
Xinwen Shu, ...

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Abstract: The next
generation of
spectroscopic surveys
will target emission-
line galaxies (ELGs)
to produce constraints

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on cosmological parameters. We study the large scale structure traced by ELGs using a combination of a semi-analytical model of galaxy formation, a code that computes the nebular emission from HII regions using the properties of the interstellar medium, and a large-volume,

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high-resolution N-
body simulation.

*[2010.08500] The
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This causes line emission spectra to be produced, as shown below. This is not a continuous spectrum as only light of specific frequencies and specific colours are produced.

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luminous (to)
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sample provides us with the most comprehensive view of reionization-era quasars to date, allowing us to connect the quasar properties with those of its host galaxy.

[2010.06902] The X-SHOOTER/ALMA sample of Quasars in

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lines occur in gases
that are optically thin,

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or in atmospheres
that have a
temperature inversion
(i.e. are hotter closer
to the observer). In an
optically thin gas, if
excited atoms or ions
are radiatively
deexcited, then the
photons produced can
escape to an
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